



## Lecture # 1

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#### Introduction

Particle physics — ultimate constituents of matter and the fundamental interactions Experiments have revealed whole families of short-lived particles Molecular hypothesis and the development of chemistry. Most scientist accepted matter aggregates of atoms.





- Radioactivity and the analysis of low energy scattering <u>atoms</u> have structure.
- Mass was concentrated in dense nucleus surrounded by cloud of electrons.
- The discovery of neutron 1930
- Geiger tubes and cloud chambers properties of cosmic ray particles.
- The modern discipline of particle physics high energy nuclear physics + cosmic ray physics





#### **Particles and Interactions**

Four interactions and their approximated strength at 10<sup>-18</sup> cm are

Strong = 1 Electromagnetic =  $10^{-2}$ Weak =  $10^{-5}$ Gravitational =  $10^{-39}$ 





- Hundreds of new particles have been discovered
- Tried to group them into families with similar characteristics.
- Leptons do not obey strong interaction.
- Hadrons obey strong interactions.
- Hadrons are of two types:
- Baryons ½ integral spin,
- Mesons —— integral spin





- Protons
- Neutrons
- Prof. Salam's weak neutral currents
- Bubble chamber
- Resonances
- Resonances can decay via strong interactions and thus have lifetime of 10<sup>-23</sup> sec
- Antimatter
- Gauge bosons





#### Detectors

- Piece of equipment for discovering the presence of something, such as metal, smoke etc
- How particles interact with matter ?
- The properties of the detectors used to measure these interactions
- Fundamental considerations involved in designing a particle physics experiment.





Charge
Mass
Spin
Magnetic moment
Life time
Branching ratios





Tracking Momentum analysis Neutral particle detection Particle identification Triggering Data acquisition





#### Alpha decay

- Radioactive decay
- Particle trapped in a potential well by nucleus
- Fundamentally quantum tunneling process
- Transition between nucleus levels
- A 5 MeV α-particle travels at 10<sup>7</sup> m/s
- Short range, 3-4 cm in air

 $Z,A) \rightarrow (Z-2,A-4)$ 







#### Beta decay

- Radioactive decay
- Fast electrons
- Weak interaction decay of neutron or proton
- Continuous energy spectrum, ranges from few keV to few tens of MeV









#### **Electron capture**

- β+ decay cannot occur in isolation
- Proton rich nuclei may also transform themselves via capture of an electron from one of the atomic orbitals
- Accompanied by electron capture process

 $energy + p + e^- \rightarrow n + \upsilon$ 

 Leaves hole, another atomic electron fills
 Emission of characteristic x-ray or auger electrons





#### Auger Electrons

- An excitation in the electron shell transferred — atomic electron rather than to a characteristic x-ray
- This occurs after electron-capture
- Second ejected electron Auger electron
- Monoenergetic energy spectrum
- Energy not more than few keV
- Susceptible to self-absorption





#### Gamma Emission

- Nucleus has discrete energy levels
- Transition between these levels by electromagnetic radiations
- Photon energy ranges keV-MeV
- Characterize high binding energy
- Y rays





#### **Annihilation Radiation**

- Annihilation of positrons
- <sup>22</sup>Na irradiate absorbing material
- Positron will annihilate with the absorber electron to produce two photons
- Photons op

opposite direction



Fig. 1.4. Gamma-ray spectrum of a <sup>22</sup>Na source as observed with a NaI detector. Because of positron annihilation in the detector and the source itself, a peak at 511 keV is observed corresponding to the detection of one of the annihilation photons





#### Internal Conversion

- Nuclear excitation energy is directly transferred to an atomic electron rather than emitting a photon
- Electron K.E = excitation energy atomic B.E
- Electrons monoenergetic
- Same energy as  $\gamma$  rays
- Few hundered keV to few MeV
- Mostly k-shell electrons ejected
- Nuclear source of monoenergetic electrons
- Used for calibration purpose





### Scattering Cross section

**Differential cross-section** 

- Gives a measure of probability for a reaction to occur
- Calculated in the form of basic interaction between the particles,



 $d\sigma$  $\frac{1}{F} \frac{dN}{d\Omega}$  $(E, \Omega)$  $\sigma(E) = \int d\Omega \, \frac{d\sigma}{d\Omega}$ 

#### **Total cross-section**

Fig. 2.1. Definition of the scattering cross section

# Energy loss by atomic collisions

Two principal features — passage of charged particle through matter
1- a loss of energy by particle
2- a deflection of the particle from its incident direction.





results of two These effects processes Inelastic collisions — atomic electrons Elastic scattering from nuclei Other process — Cherenkov radiation, nuclear reaction bremsstrahlung





- Inelastic collisions almost solely responsible
   In these collisions (ō = 10<sup>-17</sup> 10<sup>-16</sup> cm<sup>2</sup>), energy is transferred particle to the atom causing an ionization or excitation
- The amount transferred in each collision is very small fraction of the particle K.E
- Large number of collisions per unit path length
- Substantial cumulative energy loss is observed.





- Soft collisions excitation
- Hard collisions ionization
- $\delta$  -rays or knock-on electrons
- Inelastic collisions statistical in nature, their number per macroscopic path length large
- Elastic scattering from nuclei not as often as atomic collisions
- Average energy loss per unit path length
- Stopping power or  $\frac{dE}{dx}$





#### Bohr formula – Classical case Heavy particle with charge ze,M and v Calculations — impact parameter Electron is free and at initially at rest Inicident particle — undeviated Bohr formula good for heavy particles Breaks for light particles, because of quantum effects → not contain electronic coll. loss



 $-\frac{dE}{dx} = \frac{4\pi z^2 e^4}{m_e \upsilon^2} N_e \ln \frac{\gamma^2 m \upsilon^3}{z e^2 \upsilon}$ 





### The Bethe-Bloch Formula

- The energy transfer is parameterized in terms of momentum transfer rather than impact parameter.
- Momentum transfer is measureable quantity
- Impact parameter is not measureable

Shell correction

 $-\frac{dE}{dx} = 2\pi N_a r_e^2 m_e c^2 \rho \frac{Z}{A} \frac{z^2}{\beta^2}$  $\left[\ln\left(\frac{2m_e\gamma^2\upsilon^2W_{\text{max}}}{I^2}\right) - 2\beta^2 - \delta - 2$ 





 $r_e$ : Classical electron radius **Electron mass**  $N_{T}^{N_{a}}$  Avogadro's number Mean excitation potential 7 Atomic number of absorbing material Atomic weight of absorbing material Density of absorbing material Charge of incident particle v/c of incident particle **Density correction**  $\delta$ Shell correction <sup>Wmax</sup> Maximum energy transfer in one collision

#### Density effect

- Electric field of particle polarize atoms
- Electrons far from particle electric field intensity

- shielded from full
- Collisions with these outer —-contribute less total energy loss than predicted
- Energy increases velocity increases radius over which integration — increases
- Distant collisions contribute more
- This effect depends on density density effect

#### Shell correction

- Shell correction accounts velocity of particle comparable or smaller orbital velocity of electron
- At such energies assumption electron stationary not valid
- Bethe-Bloch formula breaks down
- The correction is generally small
- Other corrections also exist











 $\beta^2$ 

# Energy dependence of $\frac{dE}{dx}$

- At non-relativistic energies  $\frac{dE}{dx}$  is dominated by
- Decreases with increase of velocity until 0.96c
- Minimum ionizing
- Below the minimum ionizing each particle exhibits its own curve
- This characteristic is used to identify the particle
- At low energy region the Bethe-bloch formula breaksdown
- Energy beyond 0.96c → <sup>1</sup>/<sub>β<sup>2</sup></sub> almost constant
   <sup>dE</sup>/<sub>dx</sub> rises → logarithmic dependence
   Relativistic rise → cancelled by density correction









#### Penetration depth





At low velocity —— comparable velocity of orbital electron •  $\frac{dE}{dx}$  reaches a maximum drops sharply again. No. of complicated effects appear Tendency of the particle pickup electrons for part of the time • Lowers — effective charge — lowers  $\frac{dE}{dx}$ Heavy particle — energy deposition per unit path length ---- less at beginning ----- more at end Bragg curve





### Channeling

- Materials spatially symmetric atomic structures.
- Particle is incident at angles less than some critical angle with respect to a symmetry axis of the crystal.
- Critical angle
- Particle a series of correlated small angle scatterings
- Slowly oscillating trajectory



Schematic diagram of scattering. Particle suffers a series of correlated scatterings



#### **Critical angle**







- How far penetrate before lose all of their energy ?
- Range depends material, particle their energy.
- How calculate range
- Beam of desired energy different thickness
- Ratio transmitted to incident
- Range-number distance curve
- Range approached ratio drops.
- The curve does not drop immediately to background level.











- The curve slopes down certain spread of thickness
- Energy loss not continuous, statistical in nature.
- Two identical particles with same initial energy will not suffer the same number of collisions.
- A measurement ensemble of identical particles, — statistical distribution of ranges centered about some mean value.
- Mean range roughly half particles absorbed



This phenomenon —— range straggling Exact range all particles absorbed Tangent to the curve — at midpoint extrapolating to zero level This value — extrapolated or practical range Mean range \_\_\_\_\_\_ s(T\_0) = \int\_0^{T\_0} \left( \frac{dE}{dx} \right)^{-1} dE \right)
 Multiple scattering ---> small ---> heavy particle Semi-empirical formula  $R(T_0) = R(T_{\min}) + \int_{T_0}^{T_0} \left(\frac{dE}{dx}\right)^{-1} dE$ 





- Collision loss
- Bremsstrahlung

$$\left(\frac{dE}{dx}\right)_{tot} = \left(\frac{dE}{dx}\right)_{coll} + \left(\frac{dE}{dx}\right)_{rad}$$

- Electron-electron bremsstrahlung
- Critical energy
- Radiation length
- Range of electrons





#### **Collision** loss

Basic mechanism of collision loss valid for electrons and positrons Bethe-Bloch formula — modification Two reasons Assumption small mass —— remains undeflected → invalid indistinguishability Allowable energy transfer term  $W_{\text{max}} = \frac{T_e}{2}$  $\frac{dE}{dx} = 2 \Pi N_{a} r_{e}^{2} m_{e} c^{2} \rho \frac{Z}{A} \frac{1}{\beta^{2}} \ln \frac{\tau^{2} (\tau + 2)}{2 (I/m_{e} c^{2})^{2}} + F(\tau) - \delta - 2 \frac{C}{Z}$ 



#### Bremsstrahlung

- Small contribution few MeV or less
- At 10's of MeV, radiation loss comparable or greater than collision loss
- Dominant energy loss mechanism ——for high energy electrons —— electromagnetic radiation
- Synchrotron radiation circular acceleration
- Bremsstrahlung motion through matter
- Bremsstrahlung cross-section inverse square of particle mass

$$\frac{dE}{dt} = \left(2 \frac{e^2}{3 c^3}\right) a^2$$

lime rate of energy loss

$$\frac{d\sigma}{dk} = 5\frac{e^2}{hc}z_1^4 z_2^2 \left(\frac{mc}{M\upsilon_1}\right)^2 \frac{r_e^2}{k} \ln \frac{M\upsilon_1^2 \gamma^2}{k}$$



### Electron-electron bremsstrahlung

- E-E bremsstrahlung atomic electrons
- Critical energy  $\longrightarrow E = E_c$
- Above this enrgy radiat. loss collision-ionization loss

#### arises from field of

- dominate
- Radiation length distance over which electron energy is reduced by 1/e due to radiation loss
- Range of electrons → different from  $c \left( \frac{dE}{dx} \right)_{rad} = \left( \frac{dE}{dx} \right) coll$

 $L_{rad} = \frac{716 \, Ag \, / \, cm^2}{Z \, (Z+1) \ln \left( 287 \, / \, \sqrt{Z} \right)}$ 

800*MeV* 







For protor

Radiation loss vs collision loss for electrons in copper





### **Multiple Coulomb Scattering**

- Charged particles repeated elastic scattering from nuclei
- Small probability

Rutherford formula

 $\frac{d\sigma}{d\Omega} = z_1^2 z_2^2 r_e^2 \frac{\binom{m_e c}{\beta p}}{4 \sin^4 \left(\frac{\theta}{2}\right)}$ 

- I sin  $(\theta_2)$  dependence  $\longrightarrow$  small angular deflections
- Small energy transfer negligible
- Resultant zigzag path
- Cumulative effect is net deflection



#### Single scattering

- Thin absorber —— small prob. of more than one coulomb scattering
- Rutherford formula —— valid
- Plural scattering
- Average number of scattering < 20</p>
- Neither simple R.F nor statistical method valid



#### Multiple scattering

Average number of scattering > 20

- Small energy loss
- Statistical method to obtain net angle deflection
- Small angle approximation —— by Moliere
- Generally valid ----- upto 30 9

 $P(\theta)d\Omega = \eta d\eta \left(2\exp(-\eta^2) + \frac{F_1(\eta)}{B} + \frac{F_2(\eta)}{B^2}\right)$ 

- Backscattring of low energy electrons
- Susceptible to large angle deflections from nuclei

Moliere polar angle

Multiple scattering of a charged particle. The scale and angle an greatly exaggerated

Fig. 2.





Backscattering of low energy electrons Probability is so high, multiply and turned around altogether Absorber Backscattering out of absorber Effect strong
Iow energy electrons Depends on incident angle High-Z material Nal Non-collimated electrons, 80 % reflected back electrons incident 100m Ratio backscattered electrons





### The interaction of neutrons

- No coulomb interaction with electron or nuclei
- These interactions are rare —— short range
  - short range
     ≃10<sup>-13</sup>cm
- Neutrons must come within  $\longrightarrow \simeq 10^{13} cm$
- Normal matter mainly empty
- Neutron very penetrating particle



- Prinipal mechanism of energy loss
- Elastic scattering from nuclei MeV range
- Inelastic scattering nucleus is left in excited state gamma emission
- Neutron must have 1 MeV for inelastic collision to occur
- Radioactive neutron capture
- Neutron capture cross-section
- Valid at low energies

 $\cong \frac{1}{\upsilon}$ 

Resonance peaks superimposed upon 1/v dependence
 Other nuclear interactions (n,p), (n,d), (n,α), eV-keV

Fission

High energy hadron shower

#### conclusions

- Role of detectors in HEP
- Image: definition of energy loss calculation
  Image: definition of energy loss calculation
- Energy dependence of  $\frac{-dE}{dx}$
- Channeling
- Range
- Energy loss of electrons and positrons
- Multiple coulomb scattering
- Interaction of neutrons

# Thanks





### The interactions of photons

- Behavior of photons (x-rays, y-rays) different from charged particles
- x-rays and γ-rays are many times more penetrating
- Much smaller cross-section relative to electron inelastic collisions
- P.E, C.S and P.P remove photons from beam
- Beam of photons is not degraded
- Photoelectric effect
- Compton scattering (including Thomson and Rayleigh scattering
- Pair production)





#### Photoelectric effect

- Absorption of photon by atomic electron
- Ejection of electron from atom
- Energy of outgoing electron

 $\mathbf{E} = h \upsilon - B \cdot E$ 

- P.E always occur on bound electrons
- Nucleus absorb recoil momentum
- Cross-section increases as k-shell energy is approached
- L-absorption, M-absorption



Photoelectric cross-section as a function of incident photon energy



#### **Compton scattering**

- Best understood process in photon interaction
- Scattering of photons on free electrons
- Compton scattered cross-section
- Average fraction of total energy contained in scattered photon
- Compton absorption cross-section
- Average energy transferred to recoil electron
- Thomson and Rayleigh scatteringCoherent scattering

$$\sigma_c = \sigma^s + \sigma^a$$











Fig. 2.22. Kinematics of Compton scattering





### Pair production







Fig. 2.25. Pair production cre section in lead

### Backup slides



Thick absorber
Very thick absorber
Thin absorber